

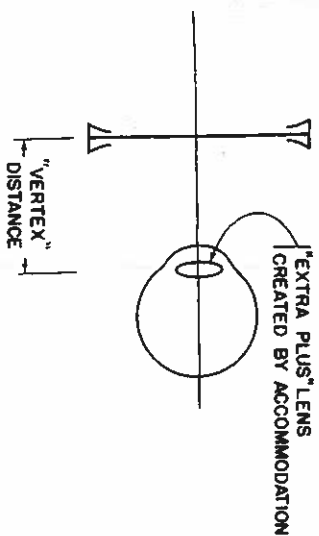
Since, in both instances, the eye is myopic 8.7 D (referred to the cornea) it is relieved of the responsibility of accommodating to the extent of the first 8.7 D. So, with the — 10 D corrective lens, this eye must accommodate (12.25 — 8.7) or 3.55 D to see the object located 20 cm from the spectacle plane; whereas the eye corrected by a contact lens must accommodate (13.35 — 8.7) or 4.65 D — that is, 1.10 D more!

(Notice from this example we assume that the contact lens-corrected myope must exert the same amount of accommodation as the emmetrope would. However, even with a contact lens, the myope will likely have to exert somewhat less. This discrepancy arises because, in our model eye, we neglected the distance behind the cornea that the myopic eye error truly "resides"; because of this oversimplification then, we find that the contact lens corrected eye is equivalent to the emmetropic one.)

#### CLINICAL POINT:

You are performing the finishing touches on a subjective refraction of a patient. If, after you arrive at his full correction, you still push on and continue to add minus lens power, the patient will tell you that the target letters seem to become smaller. Why?

The *extra* minus power you have added stimulates the patient's accommodation. That *increase* in accommodation can be looked upon as an increase in the plus power "built-in" his eye. These two dioptric powers (minus outside, plus inside) must "neutralize" each other exactly if vision is to remain clear; and so, we have created an additional "neutralized" or afocal telescopic system with the two elements separated by the existing vertex distance (plus some *more* distance within the eye to the site of the built-in plus). Since here, the "built-in" lens (which corresponds to the telescope's "eyepiece") is plus, this Galilean system is a *minifying* one for this eye.



So, placing extra minus in front of an eye forces a patient to peer through a small, reverse telescope. The greater in minus the over-correction is, the greater the accommodation required to keep the letters clear, and therefore the more the minifying effect of that telescope.

## MAGNIFICATION

So far, we have glibly tossed around the various types of magnification. We have mentioned *linear* (lateral) magnification in connection with basic lens imagery and introduced *angular* magnification when dealing with Galilean and astronomical telescopes. We have even implied the presence of a third type — called *axial*. All are often confusing, so now I want to describe each type separately and discuss it briefly so you could get the proper “feel” for all of them.

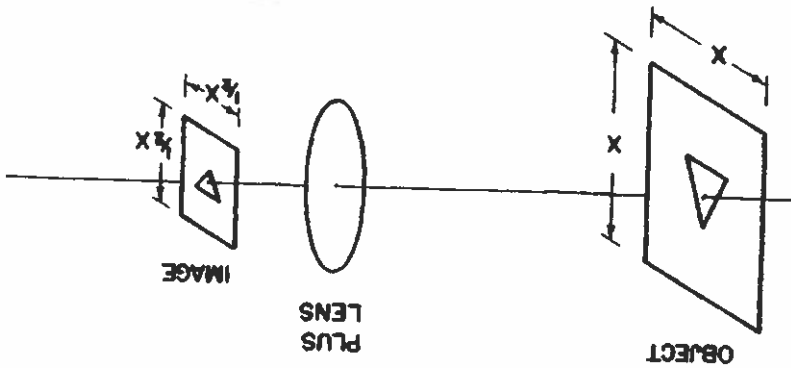
### Linear Magnification

Linear magnification has to do only with the sizes of images relative to their corresponding objects — not with how big images *look* to an eye, but how big they are. The image formed by a lens system can be larger, smaller or the same size, and it can be erect or inverted when compared to the original object. As long as we deal only in the relationships between the actual sizes (meters, inches, or microns) of objects and images, *linear* magnification is the term that applies.

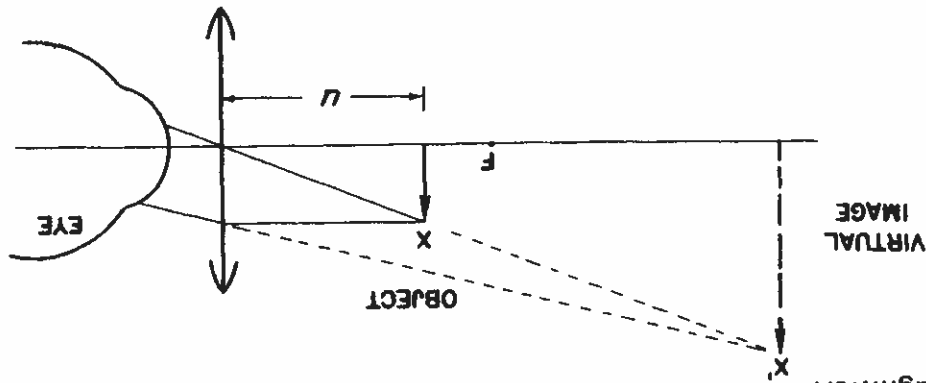
In optics and ophthalmology we need a way to deal *not* with absolute sizes, but with apparent sizes; how big or small an object or image looks to the eye. Angular size and angular magnification provide us with this means.

But *linear* magnification has no real meaning when dealing with objects at *infinite* distances. Since the linear magnification =  $\frac{\text{image size}}{\text{object size}} = \frac{\text{image distance}}{\text{object distance}}$ , if the *object* distance is infinite, the linear magnification must be zero. When the *image* distance is infinite (as when an object is placed in the focal plane of a simple plus lens), it makes no difference what the *actual* object size is, the image's size (since it is located at infinity) must be infinitely large and, therefore, the *linear* magnification is infinitely large also. In both these instances, the term *linear* (also called *transverse* as well as *lateral*) magnification has no usefulness.

LINEAR (LATERAL) MAGNIFICATION



If we know the linear magnification is  $\frac{1}{2} X$ , we know the image is half as long and half as wide as the object (the measurements taking place in the planes perpendicular to the axis of the optical system).

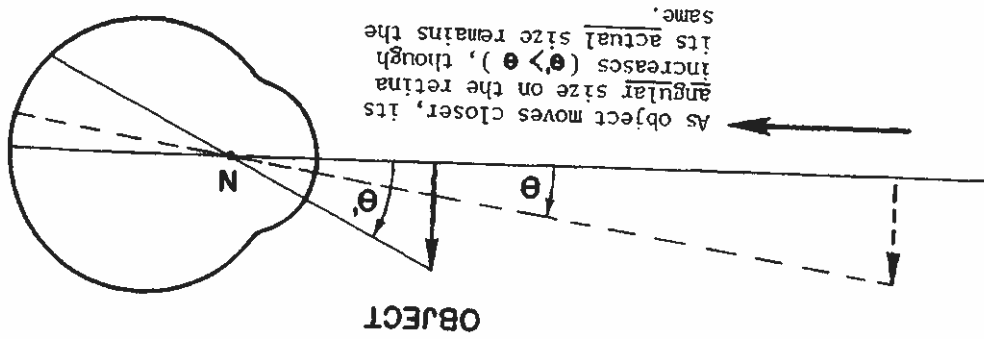


When the object is at *infinity* and an angularly magnified image is supplied by a telescope, there is really no need to specify any other fixed reference distance to calculate the angular magnification since the angular size of the *object* at infinity can serve as the reference. However, for magnifying objects close-by, we *do* need a reference distance, as you will see. Let's examine how a simple plus lens works as a magnifier.

Angular size has to do with how big an image *looks* to an eye. The same fixed size of object or image will appear larger if you approach it and smaller as you recede. The image on your retina is magnified angularly (that is, it subtends a larger angle at the eye's nodal point) as an object is brought closer. But, it is only *magnified* angularly in relation to some other angular size — that given by that same object at some fixed reference distance. For example, say an object located at an arbitrary reference distance (say 20 ft.) subtends 5° on my retina; when it approaches to half its previous distance (10 ft.), its angular size on my retina has increased to 10° — a 2 X angular magnification has occurred. If instead, the object recedes to twice its original distance (to 40 ft.), its angular size on my retina has shrunk to 2.5°, and the angular magnification has become only 1/2 X. But this is only with reference to its angular size at the 20 ft. distance. So, it should be obvious that in order to state a magnification figure, you *must* give (or assume) a reference distance; the *change* in appearance from that reference angular size is the angular magnification provided by any optical device. This point is so important and yet it is often misunderstood.

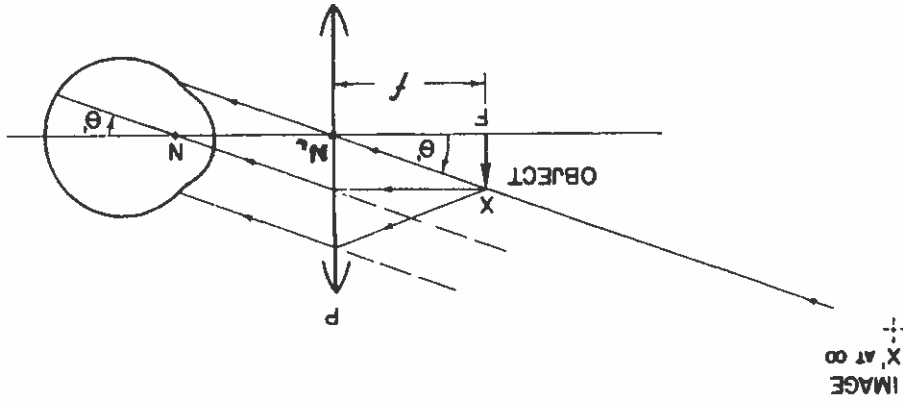
## Angular Magnification

I want particularly to emphasize the word *may* (influence the apparent size) because it *may not* also! Say the image is located at a great distance away, that is, if object distance  $u$  is equal to the focal length of the simple plus lens; now if the eye moves closer or farther from the lens (mm or yards), the *angular* size of that image on the retina will *not change*. It would only change if your eye moved an appreciable distance away from the lens as compared to the image distance, just as, for example, if you were five miles away from a mountain and moved a few meters closer to it; this would not really increase the visual size of that mountain. To accomplish that, you'd have to move say  $\frac{1}{4}$  to  $\frac{1}{2}$  mile closer. Only then might you notice any increase in its angular size on your retina. Similarly here, the image is at infinity and you would not observe any shrinkage in its angular size as you receded from it while keeping it in view through the lens.

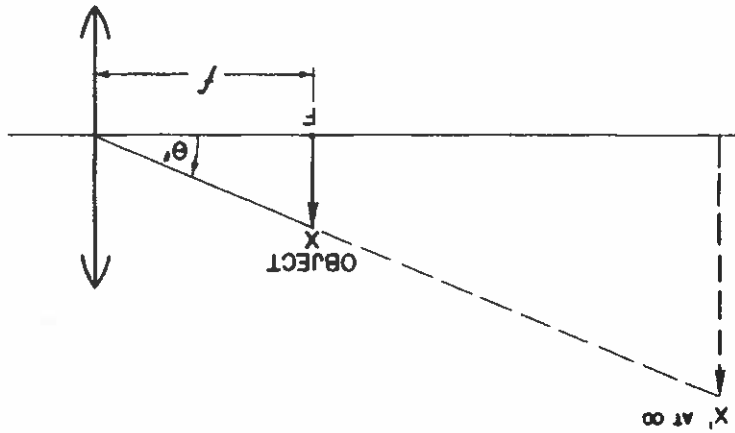


Whenever any object is placed just *inside* the primary focal point of a simple plus lens, a magnified virtual, upright image will be formed. An eye situated behind that lens will see the image. If we fix the object distance  $u$  from the lens, the image size and position will also be fixed. Any eye movement relative to the lens cannot influence the *actual* image size; however, its apparent (angular) size *may* change, just as moving closer or farther from any real life object would change its retinal image size. (As the object in the figure below moves closer to the eye, its *angular* size increases on the retina ( $\theta' > \theta$ ) though its *actual* size stays the same.)

Well, you should be able to see that *all* rays from the object tip X will, in the image space of the lens, be parallel and must all cross the axis at the same angle as that ray which goes through the lens' nodal point;



When the object is at F of the plus lens magnifier, the angular size of the *retinal* image created will also be  $\theta$ . How do we know this?



As shown in the diagram below,  $\theta'$  is the angular size of the *object* subtended at the lens;  $\theta'$  is also always the angular size of the *image* subtended at the lens.

from an eye. (See next figure.)

But as stated, the *standard* reference distance is 25 cm. An object on the visual axis would subtend angle  $\theta$  when it was placed 25 cm which might be a useful one to try first for that patient.

sized print "tells" an examiner directly the power of a low-vision aid if a given sized print can just barely be read, the M unit label for that distance. At that position, a patient with poor acuity is asked to read. example, are calibrated in M (magnification) units for a 40 cm reading. cific letter sizes of Louise Sloan's reduced-vision acuity charts, for there are certain situations where 40 cm is a better standard. The spe- here and in most other textbooks. However, you should know that years, but for most purposes 25 cm has been agreed upon and is used What that standard distance should be has been argued about for

calculated, but arbitrarily picked) as a standard.

"number". Where to put that object is what must be agreed upon (not that object anywhere you wanted and obtain some magnification pare that angular size with the  $\theta'$  created by lens P. You could put that same object, determine its angular size in that instance, and com- something. So our task is to find a reference distance at which to place tive term which states that a retinal image is bigger or smaller than having to do with the same object. Remember, magnification is a rela- must be able to compare *this* retinal image size  $\theta'$  with some *other* one But to determine how much "magnification" lens P offers us, we

$$\theta' = \frac{XF}{\text{object size}} = \frac{f}{\text{object size}} = (\text{Power of the lens}) \cdot (\text{object size})$$

considering only small angles, so  $\tan \theta' = \theta'$ ; therefore,

P of the lens, the greater  $\theta'$  will be. Remember that we are still con- proportional to  $f$ , or stated more clearly, the greater the dioptric power shorter, the retinal image size  $\theta'$  must get larger; that is,  $\theta'$  is inversely on the focal length  $f_1$  of the lens. Study the last diagram; as  $f_1$  gets

The angular size of  $\theta'$  does depend on the size of the object and which depends on the *diameter* of the lens and not its power.

which passes through lens P cannot enter the pupil — a situation And, it will always be  $\theta'$  unless the eye is so far back that light from X *eye's* nodal point, the angular size of the retinal image will be  $\theta'$  too! that angle is  $\theta'$ . Since *one* of those image rays will pass through the

eye's accommodation is in a ratio of  $\frac{4 \text{ Diopters}}{1 \text{ Mag unit}}$ , the same as that  
 With a 25 cm reference distance, the magnification given by the  
 can add to the magnification supplied by a simple plus magnifier.

it is clear that an increase in the accommodative response of an eye  
 proximity and to the required accommodation will become 2 X. Thus,  
 the retina will be doubled and the angular magnification due to this  
 accommodate the 8 D required to see it clearly, the angular size on  
 that same object is now brought closer, to 12.5 cm, and if that eye can  
 25 cm, it also would obtain a *unitary* (1 X) angular magnification. If  
 magnifiers, look at any object located at our reference distance of  
 The same reasoning applies to the eye itself. If an eye, *without* any  
 and a +20 D sphere is called a 5 X magnifier.

lens magnifier =  $\frac{4}{P}$ ; so, a +8 D sphere is called a 2 X magnifier,  
 So we have found that the angular magnification of a simple plus  
 $P \times d$  (in meters).

tance  $d$  is used as a reference, the magnification then would be  
 The  $\frac{1}{4}$  represents our chosen standard distance; if *another* dis-  
 size actually is.

and the magnification is seen to be independent of what the object

$$M = \frac{\theta}{\theta'} = \frac{\frac{\text{object size}}{.25}}{\text{object size} \cdot (P)} = \left(\frac{4}{P}\right) P$$

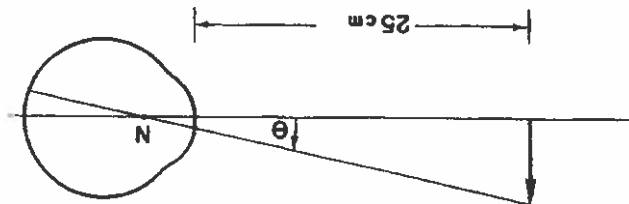
$$\theta = \frac{\text{object size}}{25 \text{ cm}} = \frac{\text{object size}}{.25 \text{ meters}}$$

and

$$\theta' = (\text{object size}) \cdot (\text{Power of the lens})$$

$P$  — the simple plus lens magnifier. With small angles:

We can return now to the angular magnification provided by lens









As shown in the above diagram, the angular image height on the retina of both the ball and the skyscraper is the same. Since a ball can come in any size, its retinal image size does not psychologically fix its distance away and the ball may be "seen" either as a small one up close or a large one farther away. A building, however, is known to be of a certain size, therefore, it is the *distance* that becomes the variable.

With both the skyscraper and the ball, I am assuming that no other cues are present which help set the distance more accurately — such as shadow, color, overlap, etc. Any of these cues (as well as your "mental set") can tip the scales towards your making a clear choice as to whether it was distance or size that *really* changed. Typically, telescopes and binoculars seem to bring objects "closer" since the objects visualized are familiar ones whose size is known. Microscopes on the other hand make objects appear "bigger" since there is no such familiarity with the object's true size. All these instruments, of course, can only enlarge the size of the retinal image of the object.

So, angular magnification (in contrast to linear magnification) involves the eye or a camera, and the true size of the image matters only indirectly, since it is only *one* of the variables, the other being its *distance*. The term "angular magnification" takes both variables into account in one datum which hides the *actual* value of either. Your personal judgment of the apparent magnitude of either will depend on the balance of many psychological factors, but for our subject "angular magnification", it is only the *angular size* that counts.

### Axial Magnification

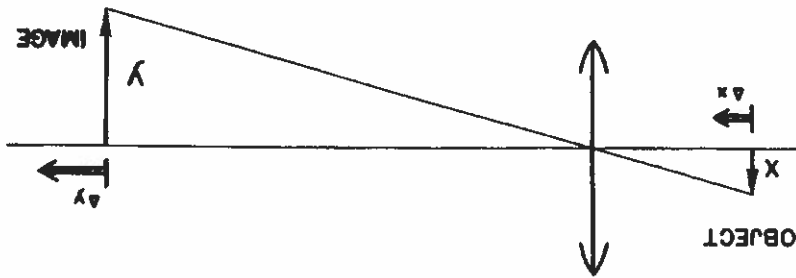
There is yet another type of magnification which spills over into both the linear and angular types; this is *axial* magnification. It is not usually discussed in the standard optics text. I am mentioning it here so that you can be aware of its presence and of its particular significance, especially in relation to indirect ophthalmoscopy.

You have been shown that with a telescope the stress on an observer's accommodation is proportional to the *square* of the *angular* magnification provided by the instrument. This is a manifestation of

This is not quite true when one deals with a refracting surface, which, of course, has different indices of refraction for the object and image spaces. For the refracting surface, the axial magnification =  $(\frac{n}{n'} \times \frac{u}{v}) \times (\frac{n}{n'}) \times (\frac{u}{v})$ . (Call this relationship "Formula One.") If you will recall, I showed (in Appendix A) that, with a refractive surface, the linear (lateral) magnification  $M$  was equal to  $(\frac{n}{n'} \times \frac{u}{v})$ . So, by substitution in "Formula One,"  $M_{axial} = (M_{linear}) \times \frac{u}{v}$ . But, with a lens in air,  $M_{linear} = \frac{u}{v}$ . So, for the lens,  $M_{axial} = (M_{linear}) \times (M_{linear})^2$ . This is what is stated in the text above.

Now you shouldn't be caught unaware when you see the term *axial* magnification again.

These shifts (axial image shift / axial object shift) are proportional to  $(\frac{x}{y})^2$ ; that is, to the square of the linear magnification. (This is detailed in one of my previous articles and will be summarized later in the section on the indirect ophthalmoscope, page 292.) So, the *axial* magnification is related to the square of the *linear* magnification "just as it is to the square of the *angular* magnification."



### AXIAL MAGNIFICATION

Even when dealing with *linear* magnification (in the figure below)  $M = \frac{\text{image size}}{\text{object size}} = \frac{y}{x}$ , axial shifts ( $\Delta x$ ) of the object will create axial shifts ( $\Delta y$ ) in the image.

and thereby demands much more accommodation. This is axial magnification. of the final image; thus the final image appears to be much closer is, as it comes closer) causes a *magnified* axial shift in the position how a small change in the object's axial (fore and aft) distance (that